## A Stage-IV CMB experiment, CMB-S4

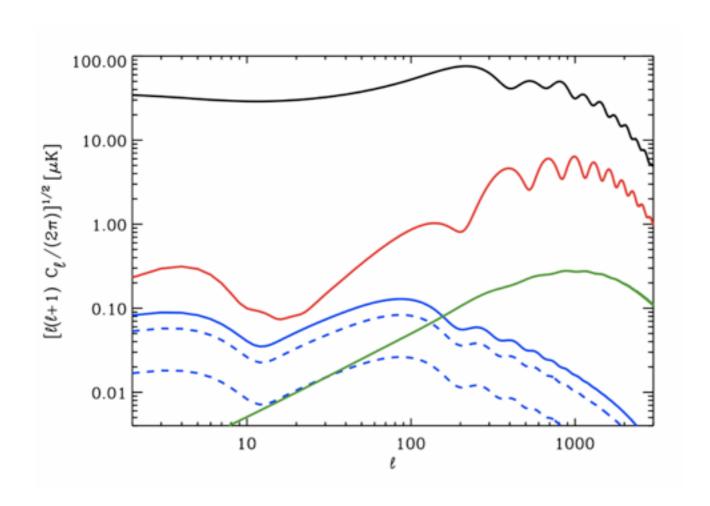
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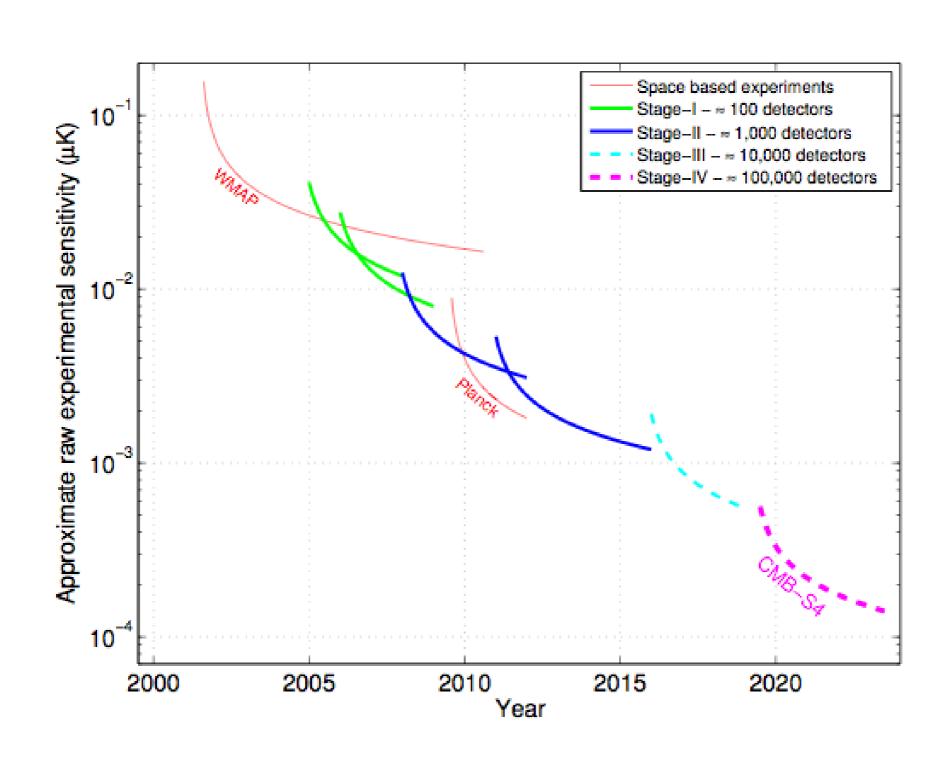
TAUP, Sept. 2013

## Goal: fully exploit CMB B-mode physics



- Explore the physics of inflation. r = gravitational wave/density wave
- Measure the Cosmic Neutrino Background.  $N_{eff}$  and  $\Sigma m_{\nu}$

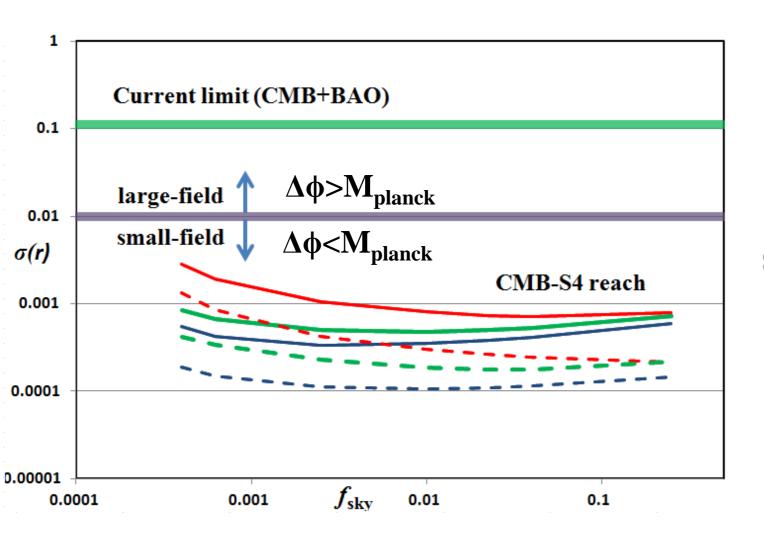
## Stages of CMB experiment

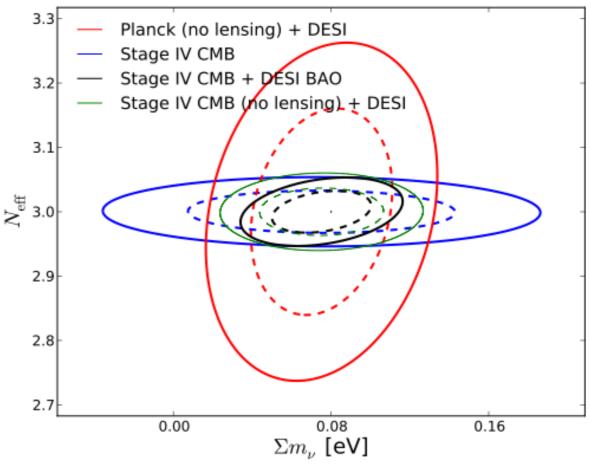


#### Science Reach of CMB-S4

Now: r < 0.73/0.12 (from *B*-modes/temperature);  $N_{eff} = 3.3 \pm 0.27$ ,  $\sum m_v < 0.23$  eV

Goal (2020+):  $\sigma(r)$ =0.001;  $\sigma(N_{eff})$ = 0.02,  $\sigma(\sum m_v)$ =0.016 eV





## Specs

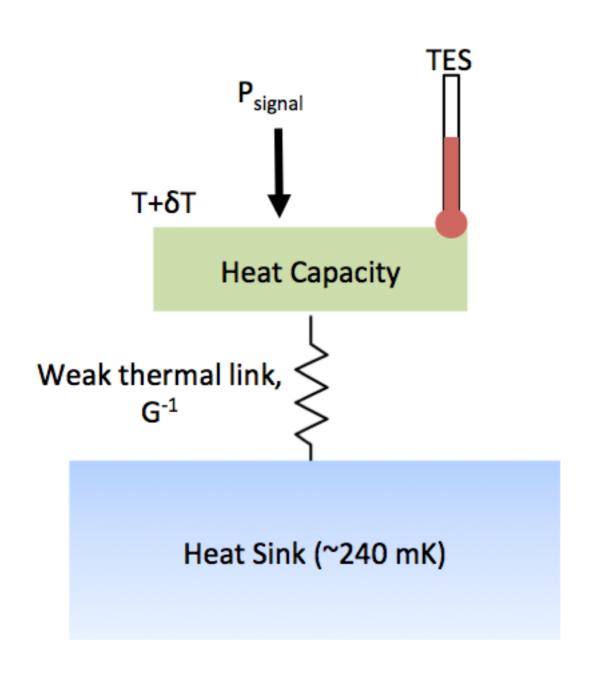
- Large angular coverage: degree angular scales for inflation, arcmin angular scales for lensing... minimum 3 arcmin resolution
- Large sky coverage: 20,000 sq deg  $(f_{sky} \sim 0.5)$
- Lots of detectors: 500,000
- Broad frequency coverage for foreground removal: 40 240 GHz
- Target noise of 1 µK-arcmin over 50% of the sky starting 2020, observing for 5 years

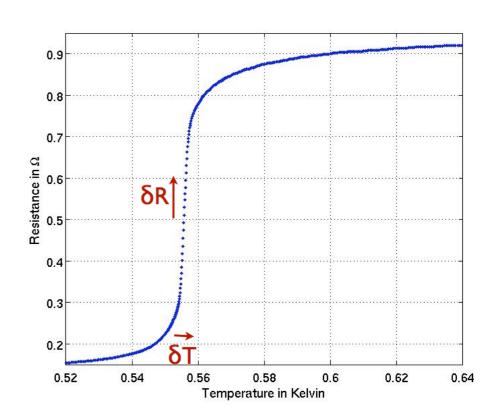
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# Primary technical development is one of scale

## Core technology: TES bolometers





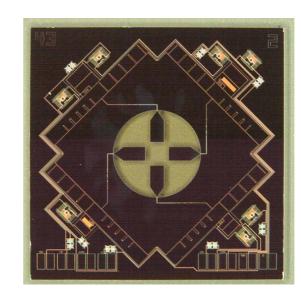
Invented by HEP for Dark Matter detection

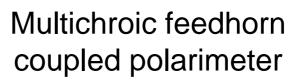
#### Core Technology: TES polarimeters





150-GHz feedhorncoupled TES polarimeters deployed at the South Pole and in the Atacama





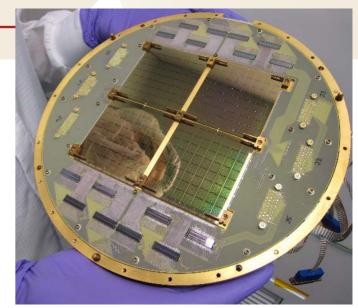


Multichroic lenscoupled polarimeter

- The scale and performance of CMB polarimeters must continue to improve to meet the needs of a stage 4 experiment.
- New direction: larger arrays of multi-frequency (multi-chroic) CMB polarimeters
- New direction: increased integration of sensor and multiplexer components.

#### Core Technology: SQUIDs

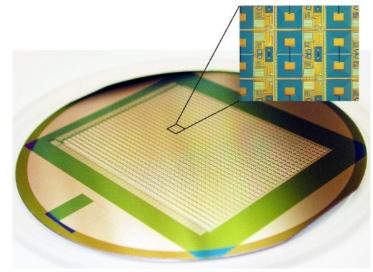
- TES arrays are doubling in size every year. We must stay on this 'Moore's Law' curve to enable a stage 4 experiment.
- This doubling requires advances in multiplexed readout electronics and Superconducting Quantum Interference Devices (SQUIDs). The maximum detectors per output channel in present CMB experiments is 32.
- Important new direction: multiplexing TESs using SQUIDs in superconducting microwave resonators enables thousands of detectors per output channel.



BICEP-2 with SQUID mux



SQUIDs for SPTpol and SPT-3G



1,280-channel SQUID multiplexer

#### Core Technology: large mm-wave components

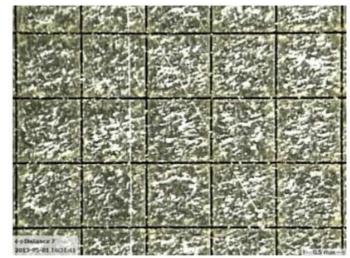
#### SLAC

## 60cm alumina lens

delivered to SLAC

for metrology





30 um cut on 1 mm grid - macro

Epoxy-based AR coating w/ grooves

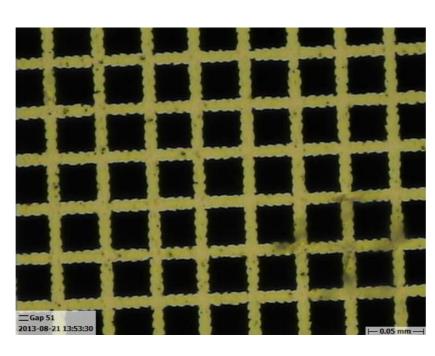
survived cryogenic cycles

## Laser machined IR filters

24" diameter (~SPT-3G window)

15μm features 3μm mylar substrate

Mm/IR properties verified





#### Technical Path for CMB-S4

- Multiplexed TES Readout: modest improvements of existing successful multiplexer technologies. New microwave-based multiplexers may provide lower cost options with broader applicability
- Large Cryogenic Optics: large aperture, large optical bandwidth cryogenic optics. Incremental improvement over Stage-III
- Improved Production Reliability: develop fabrication processes which control and stabilize material properties to achieve consistent high yield
- Increased Production Volume and Throughput: requires facilities with dedicated tooling and mass production throughput. Extensive QA testing program.
- Computing: data rate ~1 TB per day. ~10,000-times more observations per pixel compared to Planck

## Evolution of CMB experiment

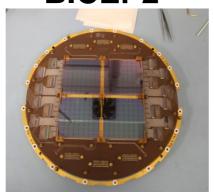
- HEP involvement in CMB has historically been small efforts focused on specific technical contributions (e.g.: multiplexing, detector fabrication and development, broadband antennas, MMICs)
- Moving from Stage-II to Stage-III and beyond, there is a consolidation of effort from multiple small experiments towards a few larger experiments
- There is increased involvement and impact from national lab resources

### CMB Focal Planes and National Labs

**BICEP** 



**BICEP2** 



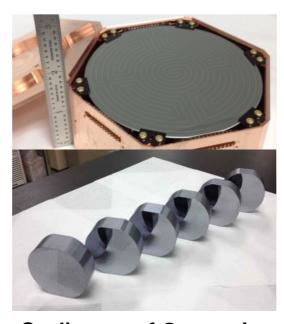
BICEP3

Fermi-LAT



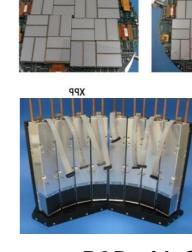
- 80 square meters of silicon sensors
- Silicon LAT assembled at SLAC

**Super CDMS** 



 Scaling up of Germanium sensors and fab throughput

**LCLS Detectors** 



- CXI-1 CXI-2
  - R&D with Cornell Univ.
  - SLAC made 10 million pixels in total so far via robotic assembly

## Summary

- CMB-S4 will fully exploit B-mode measurement. Primary challenge is one of scale.
- TES Core technology. HEP invented technology. Major impact on Stage-II incl. first lensing B-mode detection. Focus on mass production for CMB-S4.
- Change from historical *Modus Operandi*... consolidation of effort. Increased contribution/leadership from national labs.
- Impact:  $\sigma(r) = 0.001$ ;  $\sigma(N_{eff}) = 0.02$ ;  $\sigma(\Sigma m_v) = 16 \text{ meV}$